

PHY489/1489

THE DIRAC EQUATION

SO FAR:

- Quick introduction to:
 - special relativity, relativistic kinematics
 - quantum mechanics, golden rule, etc.
- Phase space:
 - how to set up and integrate over phase space to determine integrated and differential rates/cross sections
- Now we move to the particles themselves
 - start with the Dirac equation
 - describes "spin 1/2" particles
 - quarks, leptons, neutrinos

RELATIVISTIC WAVE FUNCTION

- In non-relativistic quantum mechanics, we have the Schrödinger Equation:

$$\mathbf{H}\psi = i\hbar\frac{\partial}{\partial t}\psi \quad \mathbf{H} = \frac{\mathbf{p}^2}{2m} \quad \mathbf{p} \Leftrightarrow -i\hbar\nabla$$
$$-\frac{\hbar^2}{2m}\nabla^2\psi = i\hbar\frac{\partial}{\partial t}\psi$$

- Inspired by this, Klein and Gordon (and actually Schrödinger) tried:

$$E^2 = \mathbf{p}^2 + m^2 \Rightarrow (-\nabla^2 + m^2)\psi = -\frac{\partial^2}{\partial t^2}\psi$$
$$\left(-\frac{\partial^2}{\partial t^2}\psi + \nabla^2\psi\right) = m^2\psi \quad \text{“Manifestly Lorentz Invariant”}$$

$$\partial_\mu \Rightarrow (\partial_0, \partial_1, \partial_2, \partial_3) = \left(\frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z}\right) \quad (-\partial^\mu\partial_\mu + m^2)\psi = 0$$

ISSUES WITH KLEIN-GORDON

- Within the context of quantum mechanics, this had some issues:
 - As it turns out, this allows negative probability densities: $|\psi|^2 = 0$
- Dirac traced this to the fact that we had second-order time derivative

- “factor” the E/p relation to get linear relations and obtained:

$$p_\mu p^\mu - m^2 c^2 = 0 \Rightarrow (\alpha^\kappa p_\kappa + m)(\gamma^\lambda p_\lambda - m)$$

- and found that:

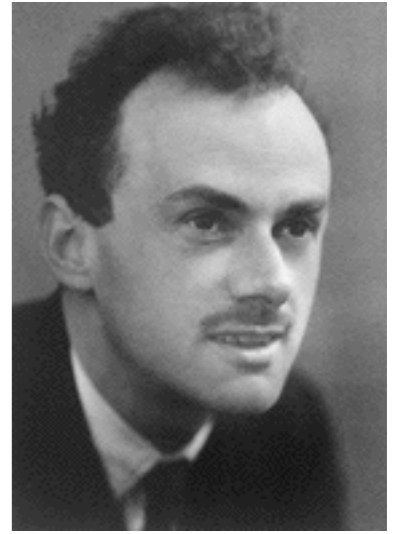
$$\alpha^\kappa = \gamma^\kappa$$

$$\{\gamma^\mu, \gamma^\nu\} = \gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2g^{\mu\nu}$$

- Dirac found that these relationships could be held by matrices, and that the corresponding wave function must be a “vector”.

$$\gamma^\mu p_\mu - m = 0 \Rightarrow (i\gamma^\mu \partial_\mu - m)\psi = 0$$

THE DIRAC EQUATION IN ITS MANY FORMS



$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad \not{\partial} \equiv a_\mu \gamma^\mu$$

$$(i\not{\partial} - m)\psi = 0 \quad \not{\partial} \equiv a_\mu \gamma^\mu = a_0 \gamma^0 - a_1 \gamma^1 - a_2 \gamma^2 - a_3 \gamma^3$$

$$\partial_\mu \Rightarrow (\partial_0, \partial_1, \partial_2, \partial_3) = \left(\frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y}, \frac{\partial}{\partial z} \right)$$

$$\left[i(\gamma^0 \partial_0 - \gamma^1 \partial_1 - \gamma^2 \partial_2 - \gamma^3 \partial_3) - m \right] \psi = 0$$

$$\left[i \left(\gamma^0 \frac{\partial}{\partial t} - \gamma^1 \frac{\partial}{\partial x} - \gamma^2 \frac{\partial}{\partial y} - \gamma^3 \frac{\partial}{\partial z} \right) - m \right] \psi = 0$$

DIRAC QUOTES:

- Q: What do you like best about America?
- A: Potatoes.

- Q: What is your favourite sport?
- A: Chinese chess.

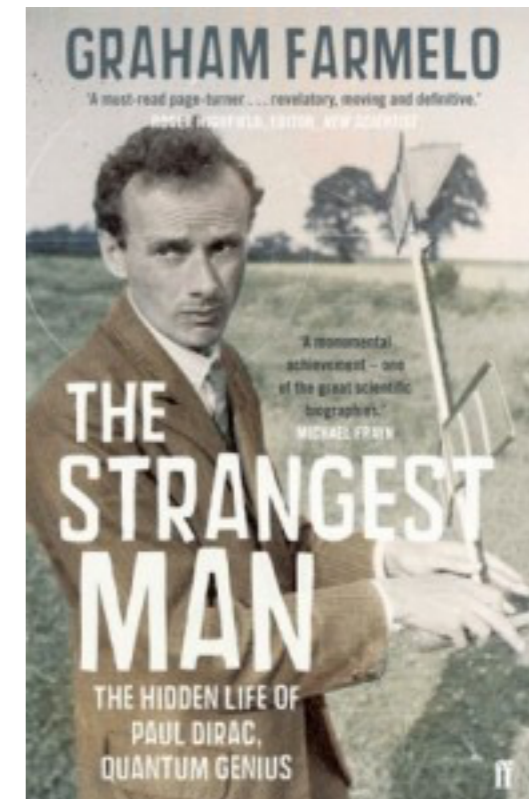
- Q: Do you go to the movies?
- A: Yes.
- Q: When?
- A: 1920, perhaps also 1930.

- Q: How did you find the Dirac Equation?
- A: I found it beautiful.

- “Dirac’s spoken vocabulary consists of “yes”, “no”, and “I don’t know.”

- Q: Professor Dirac, I did not understand how you derived the formula on the top left
- A: That is not a question. It is a statement. Next question, please.

- I was taught at school never to start a sentence without knowing the end of it.



"GAMMA" MATRICES:

$$\gamma^\mu = (\gamma^0, \gamma^1, \gamma^2, \gamma^3) \quad \vec{\sigma} = (\sigma^1, \sigma^2, \sigma^3) = \left[\begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix}, \begin{pmatrix} 0 & -i \\ i & 0 \end{pmatrix}, \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \right]$$

$$\gamma^0 = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

$$\gamma^1 = \begin{pmatrix} 0 & 0 & 0 & 1 \\ 0 & 0 & 1 & 0 \\ 0 & -1 & 0 & 0 \\ -1 & 0 & 0 & 0 \end{pmatrix} = \begin{pmatrix} 0 & \sigma^1 \\ -\sigma^1 & 0 \end{pmatrix}$$

$$\gamma^2 = \begin{pmatrix} 0 & 0 & 0 & -i \\ 0 & 0 & i & 0 \\ 0 & i & 0 & 0 \\ -i & 0 & 0 & 0 \end{pmatrix} = \begin{pmatrix} 0 & \sigma^2 \\ -\sigma^2 & 0 \end{pmatrix}$$

$$\gamma^3 = \begin{pmatrix} 0 & 0 & 1 & 0 \\ 0 & 0 & 0 & -1 \\ -1 & 0 & 0 & 0 \\ 0 & 1 & 0 & 0 \end{pmatrix} = \begin{pmatrix} 0 & \sigma^3 \\ -\sigma^3 & 0 \end{pmatrix}$$

- Note that this is a particular representation of the matrices
- Any set of matrices satisfying the anti-commutation relations works
- There are an infinite number of possibilities: this particular one (Björken-Drell) is just one example

$$\{\gamma^\mu, \gamma^\nu\} = \gamma^\mu \gamma^\nu + \gamma^\nu \gamma^\mu = 2g^{\mu\nu}$$

IN FULL GLORY

$$\left[i \begin{pmatrix} \frac{\partial}{\partial t} & 0 & -\frac{\partial}{\partial z} & -\frac{\partial}{\partial x} + i\frac{\partial}{\partial y} \\ 0 & \frac{\partial}{\partial t} & -\frac{\partial}{\partial x} - i\frac{\partial}{\partial y} & \frac{\partial}{\partial z} \\ \frac{\partial}{\partial z} & \frac{\partial}{\partial x} - i\frac{\partial}{\partial y} & \frac{\partial}{\partial t} & 0 \\ \frac{\partial}{\partial x} + i\frac{\partial}{\partial y} & -\frac{\partial}{\partial z} & 0 & -\frac{\partial}{\partial t} \end{pmatrix} - \begin{pmatrix} m & 0 & 0 & 0 \\ 0 & m & 0 & 0 \\ 0 & 0 & m & 0 \\ 0 & 0 & 0 & m \end{pmatrix} \right] \begin{pmatrix} \psi_A \\ \psi_1 \\ \psi_2 \\ \psi_3 \\ \psi_4 \\ \psi_B \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

$$p_\mu \Leftrightarrow -\partial_\mu$$

$$\begin{pmatrix} p_0 - m & -\mathbf{p} \cdot \boldsymbol{\sigma} \\ \mathbf{p} \cdot \boldsymbol{\sigma} & -p_0 - m \end{pmatrix} \begin{pmatrix} \psi_A \\ \psi_B \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$$

$$\begin{pmatrix} p_0 + m & -\mathbf{p} \cdot \boldsymbol{\sigma} \\ \mathbf{p} \cdot \boldsymbol{\sigma} & -p_0 + m \end{pmatrix} \begin{pmatrix} p_0 - m & -\mathbf{p} \cdot \boldsymbol{\sigma} \\ \mathbf{p} \cdot \boldsymbol{\sigma} & -p_0 - m \end{pmatrix} = \begin{pmatrix} p_0^2 - m^2 - (\mathbf{p} \cdot \boldsymbol{\sigma})^2 & 0 \\ 0 & p_0^2 - m^2 - (\mathbf{p} \cdot \boldsymbol{\sigma})^2 \end{pmatrix}$$

$$(\boldsymbol{\sigma} \cdot \mathbf{a})(\boldsymbol{\sigma} \cdot \mathbf{b}) = \mathbf{a} \cdot \mathbf{b} + i\boldsymbol{\sigma} \cdot (\mathbf{a} \times \mathbf{b}) \qquad (\boldsymbol{\sigma} \cdot \mathbf{p})^2 = \mathbf{p} \cdot \mathbf{p}$$

$$\begin{pmatrix} p_0^2 - m^2 - \mathbf{p}^2 & 0 \\ 0 & p_0^2 - m^2 - \mathbf{p}^2 \end{pmatrix} \begin{pmatrix} \psi_A \\ \psi_B \end{pmatrix} = \begin{pmatrix} 0 \\ 0 \end{pmatrix}$$

- This is just the KG equation four times
- Wavefunctions that satisfy the Dirac equation also satisfy KG

SOLUTIONS TO THE DIRAC EQUATION

- Particle at rest
 - General plane wave solution

$$\psi(x) \sim e^{-ik \cdot x}$$

- A particle at rest has only time dependence.

$$(i\gamma^0 \frac{\partial}{\partial t} - mc)\psi = 0 \quad \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix} \begin{pmatrix} \frac{\partial}{\partial t} \psi_A \\ \frac{\partial}{\partial t} \psi_B \end{pmatrix} = -im \begin{pmatrix} \frac{\partial}{\partial t} \psi_A \\ \frac{\partial}{\partial t} \psi_B \end{pmatrix}$$

- The equation breaks up into two independent parts:

$$\frac{\partial}{\partial t} \psi_A = -im\psi_A \quad -\frac{\partial}{\partial t} \psi_B = -im\psi_B$$

$$\psi_A = e^{-imt} \psi_A(0) \quad \psi_B = e^{+imt} \psi_B(0)$$

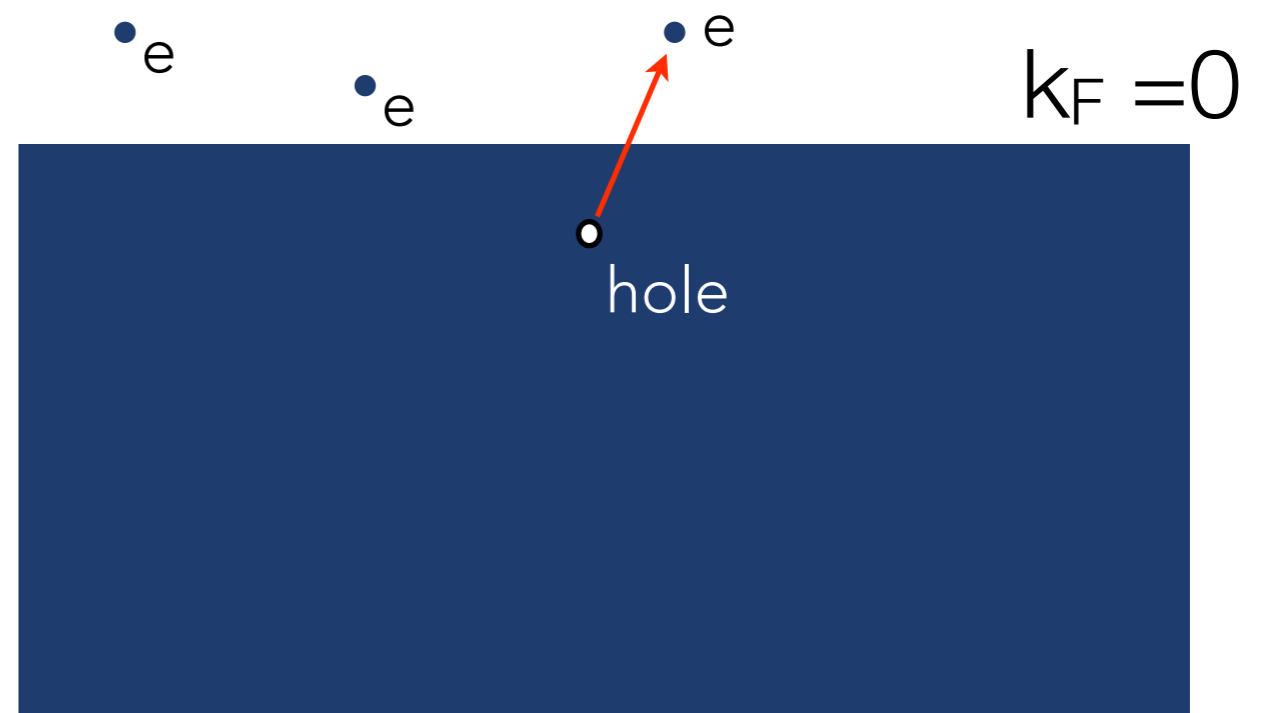
DIRAC'S DILEMMA

- ψ_B appears to have negative energy

$$\psi_A = e^{-imt} \psi_A(0) \quad \psi_B = e^{+imt} \psi_B(0)$$

- Why don't all particles fall down into these states (and down to $-\infty$)?
- Dirac's excuse: all states in the universe up to a certain level (say $E=0$) are filled.
- Pauli exclusion prevents collapse of states down to $E = -\infty$
- We can "excite" particles out of the sea into free states

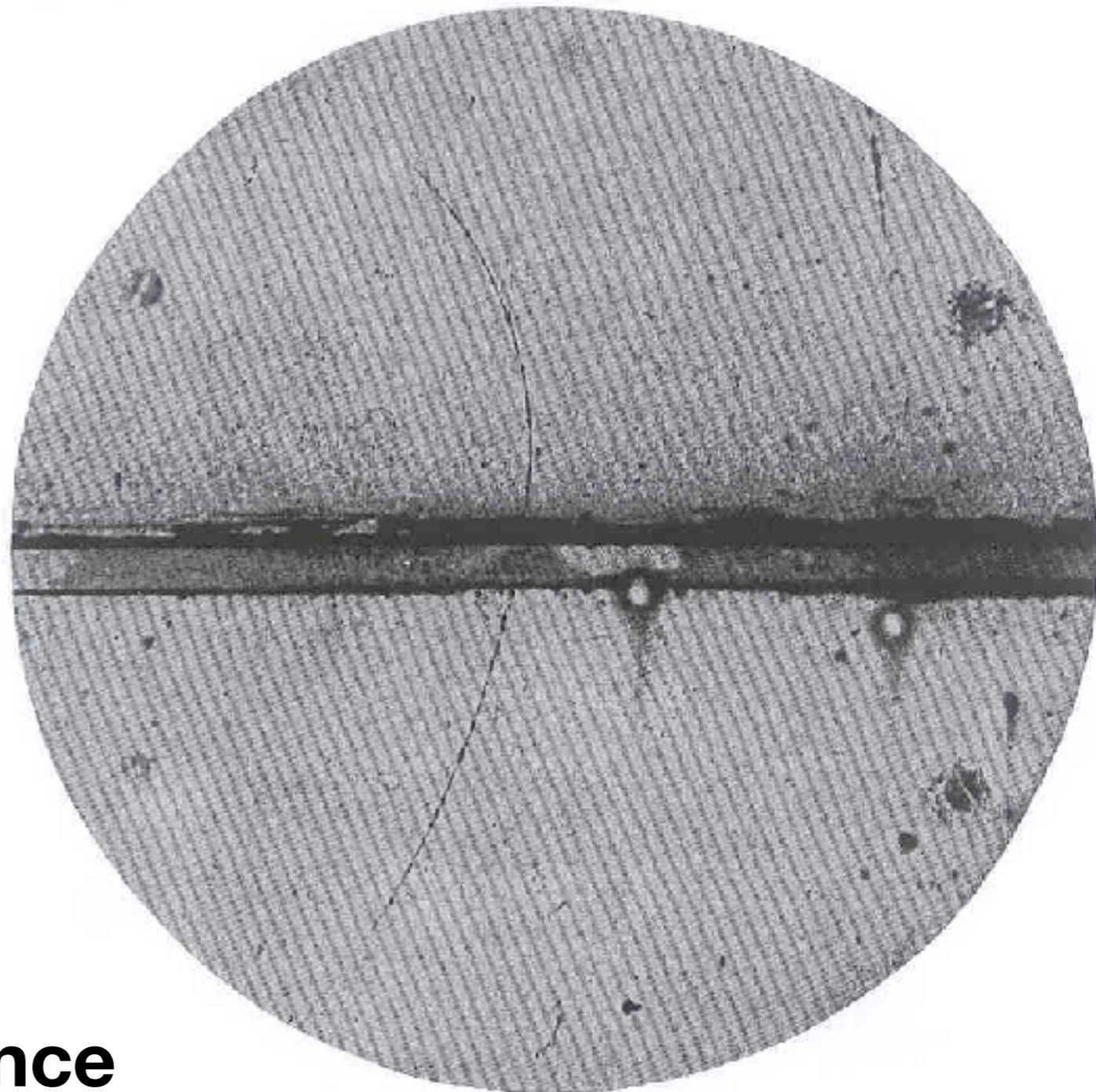
This leaves a "hole" that looks like a particle with opposite properties (positive charge, opposite spin, etc.)



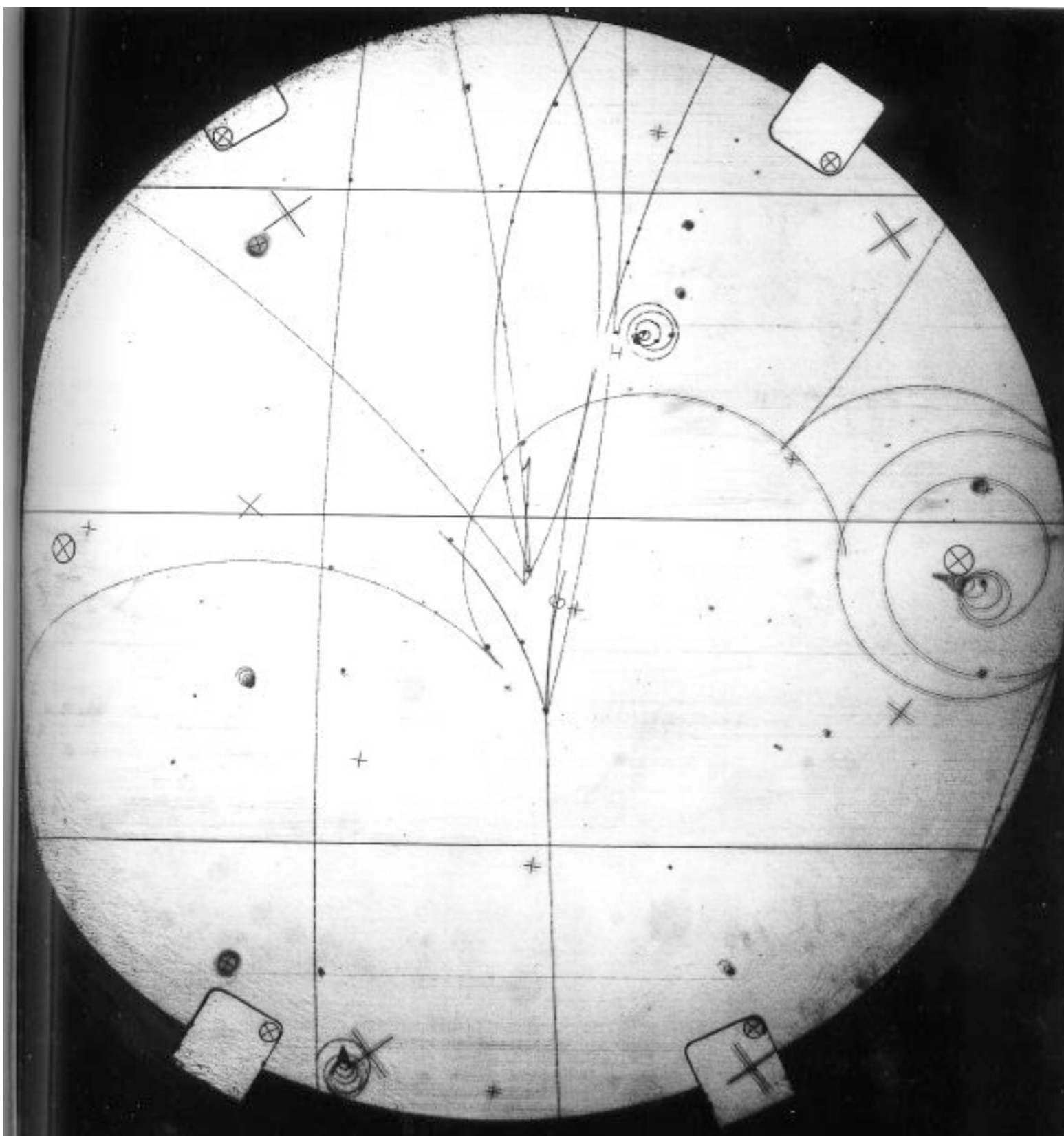
Dirac originally proposed that this might be the proton

EXCUSE TO TRIUMF

- 1932: Anderson finds "positrons"
in cosmic rays
- Exactly like electrons but
positively charged:
Fits what Dirac was looking for



**Dirac predicts the existence
of anti-matter and it is found**



SOLUTIONS TO THE DIRAC EQUATION

- Note that all particles have the same mass

$$\psi_1(t) = e^{-imc^2t/\hbar} \begin{pmatrix} 1 \\ 0 \\ 0 \\ 0 \end{pmatrix}$$

“spin up”

$$\psi_2(t) = e^{-imt} \begin{pmatrix} 0 \\ 1 \\ 0 \\ 0 \end{pmatrix}$$

“spin down”

positive energy solutions (particle)

$$\psi_3(t) = e^{+imt} \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}$$

“spin down”

$$\psi_4(t) = e^{+imt} \begin{pmatrix} 0 \\ 0 \\ 1 \\ 0 \end{pmatrix}$$

“spin up”

“negative” energy solutions (anti-particle)

PEDAGOGICAL SORE POINT

- Discussion we had thus far about difficulties with relativistic equations, (negative probabilities, negative energies) is historic
- The framework for dealing with quantum mechanics and special relativity (i.e. quantum field theory) had not been developed
 - The old tools of NR quantum mechanics had reached their limit and new ones were necessary.
 - In particular, the idea of a “wavefunction” had to be revisited
 - Until this was done, there were many difficulties!
 - Once QFT was developed, all of these problems go away.
- Both KG and Dirac Equations are valid in QFT
 - No negative probabilities, no negative energies

PLANE WAVE SOLUTIONS TO THE DIRAC EQUATION

- Consider a solution of the form:

$$\psi(x) = e^{-ik \cdot x} u(k) \quad \leftarrow \text{column vector}$$

↑
Column vector of 4 elements with space-time dependence

↑
space-time dependence

- and place it in the Dirac equation:

$$(i\gamma^\mu \partial_\mu - m)\psi = 0$$

$$(\gamma^\mu k_\mu - m)e^{-ik \cdot x} u(k)\psi = 0 \quad (\gamma^\mu k_\mu - m)u(k)\psi = 0$$

MORE EXPLICITLY:

- By definition:

$$\gamma^\mu k_\mu = \gamma^0 k^0 - \gamma^1 k^1 - \gamma^2 k^2 - \gamma^3 k^3 \quad \longrightarrow \quad \begin{pmatrix} k_0 & -\mathbf{k} \cdot \boldsymbol{\sigma} \\ \mathbf{k} \cdot \boldsymbol{\sigma} & -k_0 \end{pmatrix}$$

- So that the Dirac equation reads:

$$(\gamma^\mu k_\mu - m)u(k) = 0$$

Note 2x2 notation

$$u(k) \Rightarrow \begin{pmatrix} u_A \\ u_B \end{pmatrix}$$

$$\begin{pmatrix} k_0 - mc & -\mathbf{k} \cdot \boldsymbol{\sigma} \\ \mathbf{k} \cdot \boldsymbol{\sigma} & -k_0 - mc \end{pmatrix} \begin{pmatrix} u_A \\ u_B \end{pmatrix}$$

$$(k_0 - m)u_A - (\mathbf{k} \cdot \boldsymbol{\sigma})u_B = 0$$

$$u_A = \frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 - m} u_B$$

$$(\mathbf{k} \cdot \boldsymbol{\sigma})u_A - (k_0 + m)u_B = 0$$

$$u_B = \frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 + m} u_A$$

DETERMINING u

- this means we can identify $k \leftrightarrow \pm p$

$$u_A = \frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 - m} u_B \quad u_B = \frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 + m} u_A$$

$$\frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 + m} \frac{\mathbf{k} \cdot \boldsymbol{\sigma}}{k_0 - m} = 1$$

$$p^\mu = \pm k^\mu$$

$$\mathbf{p} \cdot \boldsymbol{\sigma} = \begin{pmatrix} p_z & p_x - ip_y \\ p_x + ip_y & -p_z \end{pmatrix}$$

- We can now construct the column vector u :

electrons

$$u_1 = N \begin{pmatrix} 1 \\ 0 \\ p_z/(E+m) \\ (p_x + ip_y)/(E+m) \end{pmatrix} \quad u_2 = N \begin{pmatrix} 0 \\ 1 \\ (p_x - ip_y)/(E+m) \\ -p_z/(E+m) \end{pmatrix}$$

Used positive k solutions

positrons

$$u_3 = N \begin{pmatrix} p_z/(E+m) \\ (p_x + ip_y)/(E+m) \\ 1 \\ 0 \end{pmatrix} \quad u_4 = N \begin{pmatrix} (p_x - ip_y)/(E+m) \\ -p_z/(E+m) \\ 0 \\ 1 \end{pmatrix}$$

Used negative k solutions

NORMALIZATION:

- Choose “normalization” of the wavefunctions
 - Note that multiples of the solutions are still solution
 - normalization convention simply fixes this arbitrary choice:

$$u^\dagger u = 2E \quad u^\dagger \equiv (u^T)^* \quad u = \begin{pmatrix} u_1 \\ u_2 \\ u_3 \\ u_4 \end{pmatrix} \Rightarrow u^\dagger = (u_1^*, u_2^*, u_3^*, u_4^*)$$

- for u_1

$$u_1^\dagger u_1 = N^2 \left[1 + \frac{\mathbf{p}^2}{(E + m)^2} \right] = 2E \quad N = \sqrt{E + m}$$

LORENTZ COVARIANCE

- The Dirac equation “works” in all reference frames.
 - What exactly does this mean?

$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad \text{“Lorentz Covariant”}$$

- i , m , γ are constants that don’t change with reference frames.
- ∂_μ and ψ will change with reference frames, however.
 - ∂_μ is a derivative that will be taken with respect to the space-time coordinates in the new reference frame. We’ll call this ∂'_μ
 - how does ψ change?
 - $\psi' = S\psi$ where ψ' is the spinor in the new reference frame

- Putting this together, we have the following transformation of the equation when evaluating it in a new reference frame

$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad \Rightarrow \quad (i\gamma^\mu \partial'_\mu - m)\psi' = 0$$

$$i\gamma^\mu \partial'_\mu (S\psi) - m(S\psi) = 0$$

PROPERTIES OF S

- In general, we know how to relate space time coordinates in one reference with another (i.e. Lorentz transformation), we can do the same for the derivatives
- Using the chain rule, we get: $\partial'_{\mu} \equiv \frac{\partial}{\partial x^{\mu'}} = \frac{\partial x^{\nu}}{\partial x^{\mu'}} \frac{\partial}{\partial x^{\nu}}$
 - where we view x as a function x' (i.e. the original coordinates as a function of the transformed or primed coordinates).
 - Note the summation over ν
- if the primed coordinates moving along the x axis with velocity β :

$$\begin{aligned} x^0 &= \gamma(x^{0'} + \beta x^{1'}) & (\nu = 0, \mu = 0) &\Rightarrow \frac{\partial x^0}{\partial x^{0'}} = \gamma \\ x^1 &= \gamma(x^{1'} + \beta x^{0'}) \\ x^2 &= x^{2'} \\ x^3 &= x^{3'} & (\nu = 0, \mu = 1) &\Rightarrow \frac{\partial x^0}{\partial x^{1'}} = \gamma\beta \end{aligned}$$

TRANSFORMING THE DIRAC EQUATION

$$i\gamma^\mu \partial_\mu \psi - m\psi = 0 \quad \Rightarrow \quad i\gamma^\mu \partial'_\mu \psi' - m\psi' = 0$$

$$i\gamma^\mu \partial'_\mu (S\psi) - m(S\psi) = 0$$

$$i\gamma^\mu \frac{\partial x^\nu}{\partial x^{\mu'}} \partial_\nu (S\psi) - m(S\psi) = 0$$

S is constant in space time, so we can move it to the left of the derivatives

$$i\gamma^\mu S \frac{\partial x^\nu}{\partial x^{\mu'}} \partial_\nu \psi - mS\psi = 0$$

Now slap S^{-1} from the left

$$(i\gamma^\mu \partial_\mu - m)\psi = 0 \quad S^{-1} \rightarrow i\gamma^\mu S \frac{\partial x^\nu}{\partial x^{\mu'}} \partial_\nu \psi - mS\psi = 0$$

Since these equations must be the same, S must satisfy

$$\gamma^\nu = S^{-1} \gamma^\mu S \frac{\partial x^\nu}{\partial x^{\mu'}}$$

PARITY OPERATOR

- For the parity operator, invert the spatial coordinates while keeping the time coordinate unchanged:

$$P = \begin{pmatrix} 1 & 0 & 0 & 0 \\ 0 & -1 & 0 & 0 \\ 0 & 0 & -1 & 0 \\ 0 & 0 & 0 & -1 \end{pmatrix} \quad \frac{\partial x_0}{\partial x_0'} = 1 \quad \frac{\partial x_1}{\partial x_1'} = -1$$

$$\frac{\partial x_2}{\partial x_2'} = -1 \quad \frac{\partial x_3}{\partial x_3'} = -1$$

- We then have

$$\gamma^0 = S^{-1} \gamma^0 S$$

$$\gamma^1 = -S^{-1} \gamma^1 S$$

$$\gamma^2 = -S^{-1} \gamma^2 S$$

$$\gamma^3 = -S^{-1} \gamma^3 S$$

$$\gamma^\nu = S^{-1} \gamma^\mu S \frac{\partial x^\nu}{\partial x^{\mu'}}$$

$$\gamma^0 = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$

$$(\gamma^0)^2 = 1$$

$$\{\gamma^\mu, \gamma^\nu\} = 2g^{\mu\nu} \Rightarrow \gamma^0 \gamma^i = -\gamma^i \gamma^0$$

We find that γ^0 satisfies our needs

$$\gamma^0 = \gamma^0 \gamma^0 \gamma^0 = \gamma^0$$

$$\gamma^i = -\gamma^0 \gamma^i \gamma^0 = \gamma^0 \gamma^0 \gamma^i = \gamma^i$$

$$S_P = \gamma^0$$

NEXT TIME

- Read 4.6-4.9 and Chapter 5
- Lots of notation, lots of stuff going on
 - please stop by if you have questions!